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Multi-wavelength spin dynamics of defects in hexagonal boron nitride

Ivan Zhigulin^{1,2}, Nicholas P. Sloane^{1,2}, Benjamin Whitefield^{1,2}, Konosuke Shimazaki^{1,2}, Jean-Philippe Tetienne³, Mehran Kianinia^{1,2} and Igor Aharonovich^{1,2,4}

Abstract

Optically addressable solid-state spin defects are essential platforms for quantum sensing and information processing. Recently, single spin defects with combined $S = 1$ and $S = \frac{1}{2}$ spin transitions were discovered in hexagonal boron nitride (hBN). In this work we unveil their excitation dynamics. In particular, we study the effects of the excitation wavelength on the spin-dependent fluorescence and the spin dynamics of these peculiar quantum spin defects. We find that changing the excitation wavelength leads to a threefold enhancement in both the optically detected magnetic resonance (ODMR) contrast and the corresponding magnetic field sensitivity. In addition, we find that the excitation wavelength has a strong impact on the photodynamics of spin complex emitters. Our work presents valuable insights to the mechanistic understanding of spin complex emitters in hBN and highlights the importance of excitation wavelength for optimising their performance in quantum sensing and quantum technologies.

Introduction

Optically addressable spin systems have emerged as promising platforms for quantum communication and sensing^{1–3}. Defects in three-dimensional host materials such as diamond and silicon carbide have been extensively studied for quantum sensing, demonstrating sensitivity to magnetic fields, electric fields, and temperature⁴. Alongside these conventional platforms, hexagonal boron nitride (hBN) has emerged as a complimentary host for optically addressable spin defects, with its van der Waals structure providing unique advantages^{5–13}. Notably, the two-dimensional nature of hBN allows atomically thin layers to be exfoliated directly onto target samples^{14–18} or seamlessly integrated into photonic structures and optoelectronic devices^{19–21}.

Currently, the most studied spin defect in hBN is the boron vacancy (V_B^-) which has a spin-triplet ($S = 1$) ground state^{22–26}. This defect exhibits optically detected

magnetic resonance (ODMR) at zero field and at room temperature. It has already been utilised for sensing magnetic fields, temperature, pressure, and strain^{27–30}. However, so far single V_B^- defects have not been identified, which limits its broad adaptation for quantum information. In addition, their low quantum efficiency and the dark zero phonon line has further motivated a search for bright narrowband single spin defects in hBN.

Recently, a new class of spin-active defects has been discovered in hBN and other solid-state materials, referred to as the spin complex^{10–12,31–33}. These spin complex defects are brighter, show a clear single photon emission, and have zero-phonon lines (ZPLs) spanning the range of visible to near-infrared wavelengths (500–750 nm)¹². These systems have also been analytically studied using ab initio theoretical approaches^{34,35}. Most intriguingly, the spin complex in hBN exhibits both $S = 1$ and $S = \frac{1}{2}$ -like spin transitions in ODMR spectra. The $S = 1$ transitions are attributed to an electron pair occupying a strongly coupled triplet state, whereas the $S = \frac{1}{2}$ transitions correspond to a delocalised, weakly coupled electron spin pair^{10,12}. The coexistence of these spin manifolds provides an exciting new platform for quantum sensing and spin manipulation in hBN.

Correspondence: Nicholas P. Sloane (Nicholas.sloane@uts.edu.au) or Mehran Kianinia (Mehran.Kianinia@uts.edu.au)

¹School of Mathematical and Physical Sciences, University of Technology Sydney, Ultimo, NSW 2007, Australia

²ARC Centre of Excellence for Transformative Meta-Optical Systems, University of Technology Sydney, Ultimo, NSW 2007, Australia

Full list of author information is available at the end of the article

These authors contributed equally: Ivan Zhigulin, Nicholas P. Sloane

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In this work, we study the spin-photon dynamics of the spin complex in hBN, focusing specifically on the excitation wavelength and its effect on the spin dynamics. We find that the ODMR contrast and the stability of the photoluminescence (PL) of the spin complex are strongly influenced by the excitation wavelength, with ODMR transitions tripling in contrast and reaching nearly 100%.

Results

The spin complex in hBN is schematically illustrated in Fig. 1a, which exhibits two distinct spin manifolds, a localised strongly coupled spin pair (orange) and a delocalised weakly coupled spin pair (blue). The energetic structure formed by the spin complex is shown in Fig. 1b and includes the optically active $S = 0$ ground and excited states ($|g\rangle$ and $|e\rangle$, respectively), as well as the triplet states in the metastable (MS) regime which consists of two distinct spin manifolds. The first manifold has both electrons occupying the same defect resulting in a localised, strongly coupled spin pair with the well defined spin states ($S = 1$). These energetic levels are shown in the red box in Fig. 1b as $|+1\rangle$, $|0\rangle$, and $|-1\rangle$ corresponding to $m_s = +1$, 0, and -1 , respectively. Following charge transfer of one electron to a nearby defect, the spin pair becomes delocalised, resulting in weaker coupling and transition into the second manifold ($S = \{1,0\}$). In this configuration the spin states are instead effectively characterised by a state of mixed singlet and triplet character ($|ST_0\rangle$) alongside states with pure triplet character ($|T_{\pm}\rangle$) as shown in Fig. 1b. From the weakly coupled manifold the electron can reform back to the optically active defect via charge transfer, relaxing to the $S = 0$ ground state $|g\rangle$, from which it can be re-excited and undergo fluorescence.

As $|ST_0\rangle$ and $|T_{\pm}\rangle$ decay at different rates to $|g\rangle$, applying microwaves resonant with the splitting between the weakly coupled spin states leads to a change in the fluorescence, resulting in a detectable ODMR signal. Similarly, before charge transfer to the weakly coupled manifold occurs, application of resonant microwaves with the splitting between the strongly coupled spin states modifies the population of $|ST_0\rangle$ and $|T_{\pm}\rangle$, resulting in an ODMR signal. Previous studies have shown that no ODMR is observed at zero field, implying that the strongly coupled spin-pair population cascades through the weakly coupled manifold prior to relaxation to the ground state^{10,12}. At low fields, the energetic splitting between the spin states in the weakly coupled manifold is nearly degenerate leading to rapid spin mixing and loss of spin information.

Figure 1c shows the typical continuous wave (CW)-ODMR spectrum from a spin complex emitter with signatures of transitions between both the strongly coupled ($S = 1$) states and the weakly coupled ($S = \{1,0\}$) states. Resonances in the ODMR spectrum are assigned to the

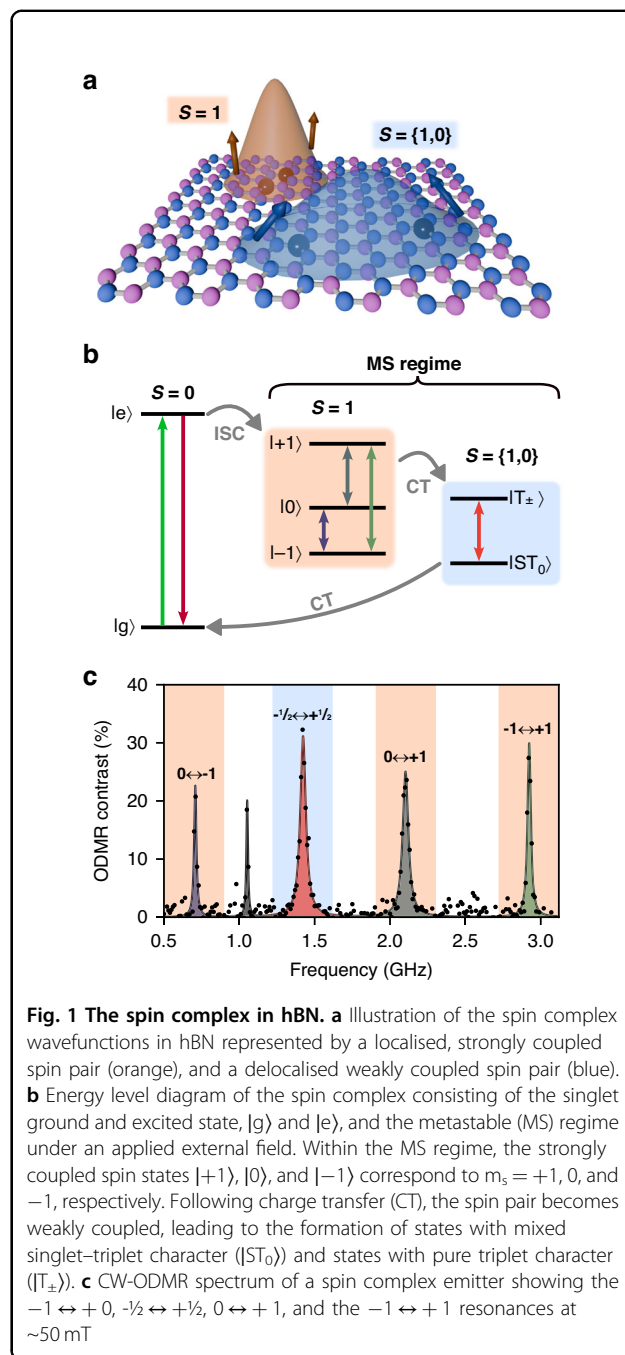


Fig. 1 The spin complex in hBN. **a** Illustration of the spin complex wavefunctions in hBN represented by a localised, strongly coupled spin pair (orange), and a delocalised weakly coupled spin pair (blue). **b** Energy level diagram of the spin complex consisting of the singlet ground and excited state, $|g\rangle$ and $|e\rangle$, and the metastable (MS) regime under an applied external field. Within the MS regime, the strongly coupled spin states $|+1\rangle$, $|0\rangle$, and $|-1\rangle$ correspond to $m_s = +1$, 0, and -1 , respectively. Following charge transfer (CT), the spin pair becomes weakly coupled, leading to the formation of states with mixed singlet-triplet character ($|ST_0\rangle$) and states with pure triplet character ($|T_{\pm}\rangle$). **c** CW-ODMR spectrum of a spin complex emitter showing the $-1 \leftrightarrow +1$, $-1/2 \leftrightarrow +1/2$, $0 \leftrightarrow +1$, and the $-1 \leftrightarrow +1$ resonances at ~ 50 mT

different transitions based on solutions to the spin Hamiltonian incorporating the $S = 1$ and $S = \{1,0\}$ states (see Supplementary Information Section 1). As no ODMR is detected at zero field for the spin complex emitter, the spectrum was acquired with an applied out-of-plane magnetic field of ~ 50 mT. The magnetic field strength was kept approximately the same for all measurements throughout this work by ensuring that the $S = 1/2$ transition occurred at approximately the same frequency. The transitions within the strongly coupled manifold, namely



$-1 \leftrightarrow 0$, $0 \leftrightarrow +1$, and the $-1 \leftrightarrow +1$ resonances can be seen at approximately 0.7 GHz, 2.1 GHz, and 2.9 GHz, respectively. The peak at ~ 1.1 GHz corresponds to the second harmonic of the $0 \leftrightarrow +1$ transition. Furthermore, we note that the $-1 \leftrightarrow +1$ transition at 2.9 GHz indicates a double-quantum transition ($\Delta m_s = 2$), which is forbidden under standard spin-selection rules. Nevertheless, this transition produces a strong ODMR signal and has been widely reported for spin complex emitters. The transition within the weakly coupled manifold ($S = \{1,0\}$) is observed at 1.4 GHz and is commonly referred to as the $-\frac{1}{2} \leftrightarrow +\frac{1}{2}$ transition, as its response is reminiscent to the behaviour of a single spin- $\frac{1}{2}$ particle.

Prior to entering the MS regime via intersystem crossing (ISC), the spin complex defect must first be driven to the excited state by an absorption of incoming photons. To locate emitters, scanning confocal microscopy was employed. Figure SI2a shows a region of hBN flake with multiple optically active emitters, highlighting the emitter of interest. PL spectra of this emitter under 532 nm and 633 nm excitations are presented in Figure SI2b, with no discernible differences observed between the two wavelengths. Similarly, Fig. SI 3a, b shows the autocorrelation measurements of the emitter, which exhibit comparable antibunching behaviour ($g^2(0) < 0.5$) under both 532 nm and 633 nm excitations. We note that this emitter could not be resolved under 405 nm excitation, likely due to a reduced absorption cross-section at this wavelength or excitation into an optically inactive state. Alternatively, higher-energy 405 nm excitation may induce charge state switching or photoionization, potentially resulting in a non-emissive charge configuration.

In stark contrast to the PL signatures of this emitter under different excitation wavelengths, the spin-dependent fluorescence is strongly impacted by the excitation wavelength. Figure 2a presents CW-ODMR spectra of the $-\frac{1}{2} \leftrightarrow +\frac{1}{2}$ and $0 \leftrightarrow +1$ transitions under 532 nm and 633 nm excitations. Remarkably, the ODMR contrast is approximately three times greater when excited with 633 nm relative to 532 nm excitation, increasing from 36 to 98% for the $-\frac{1}{2} \leftrightarrow +\frac{1}{2}$ transition. This behaviour is similarly observed for all ODMR transitions outlined in Fig. 1b, with the changes in the $-1 \leftrightarrow +0$ and $-1 \leftrightarrow +1$ transitions shown in Fig. SI4.

Figure 2b shows the relationship between the measured ODMR contrast for the $-\frac{1}{2} \leftrightarrow +\frac{1}{2}$ and $0 \leftrightarrow +1$ transitions and the excitation power (P_{exc}), normalised by the saturation power (P_{sat}) extracted from Fig. SI5 for both excitation wavelengths. The obtained saturation powers from Fig. SI5 were found to be 151 μW and 97 μW for 532 nm and 633 nm excitations, respectively. Note the rise and subsequent fall in contrast as a function of optical excitation power, which was also observed previously for the spin complex in hBN¹⁰. Yet, the contrast under

633 nm excitation consistently remains higher than that under 532 nm excitation across all laser powers, suggesting that this behaviour is not described by the optical excitation power. Instead, the enhanced contrast may potentially arise from different couplings between the excited states accessed by each wavelength and the spin-active MS states. Measurements of Rabi oscillations and readout delay are also presented and discussed in Supplementary Information Section 1 and exhibit the same behaviour as the CW-ODMR results, namely an increased ODMR contrast under 633 nm excitation compared to 532 nm excitation. We measured additional spin complex emitters and observed similar excitation wavelength-dependence of the ODMR signal (see Supplementary Information Section 1). Note that this effect appears to be emitter-dependent, with different excitation wavelengths producing varying contrast strengths across different emitters. A summary of the ODMR contrast for all characterised emitters is provided in Table SI2. This stems from the diverse defect configurations in hBN responsible for quantum emission^{36–39}, and is consistent with previous reports demonstrating no correlations between ZPL and contrast^{10,12}.

To investigate the wavelength-dependent photodynamics of the emitter, PL counts were continuously measured over an extended time interval. Figure 2c, d shows photodetector count traces of PL over time of the emitter excited with 532 nm and 633 nm, respectively. For additional comparison, PL count histograms for both excitation wavelengths are also included in Fig. 2c, d. Under 532 nm excitation, PL of the emitter remains stable for the duration of the measurement. In contrast, under 633 nm excitation, the emitter exhibits pronounced “blinking” behaviour, intermittently switching between bright emission and non-luminescent intervals as seen in the variation in the histogram in Fig. 2d. To further probe the non-trivial photodynamics, second-order autocorrelation with longer timescales (~ 1 ms) was performed under both 532 nm and 633 nm excitations at $\sim 0.5P_{\text{sat}}$ for each. These measurements, presented in Fig. 2e, show bunching signatures for both wavelengths with a stronger amplitude for the 633 nm. To elucidate the discrepancy, Table SI1 in Section 2 of the Supplementary Information examines the obtained fit parameters. These demonstrate similar radiative lifetimes ($\tau_{\text{rad}} \cong 1.75$ ns) for both wavelengths, with 633 nm excitation displaying a significantly increased population of metastable/shelving states alongside an additional long-lived component (~ 460 μs) arising from the blinking behaviour. Besides the stronger coupling to the MS regime, full description of the observed effect cannot also neglect the presence of optically inactive states that are being populated, including other excited energy charge states or electron transfers to nearby charge traps.



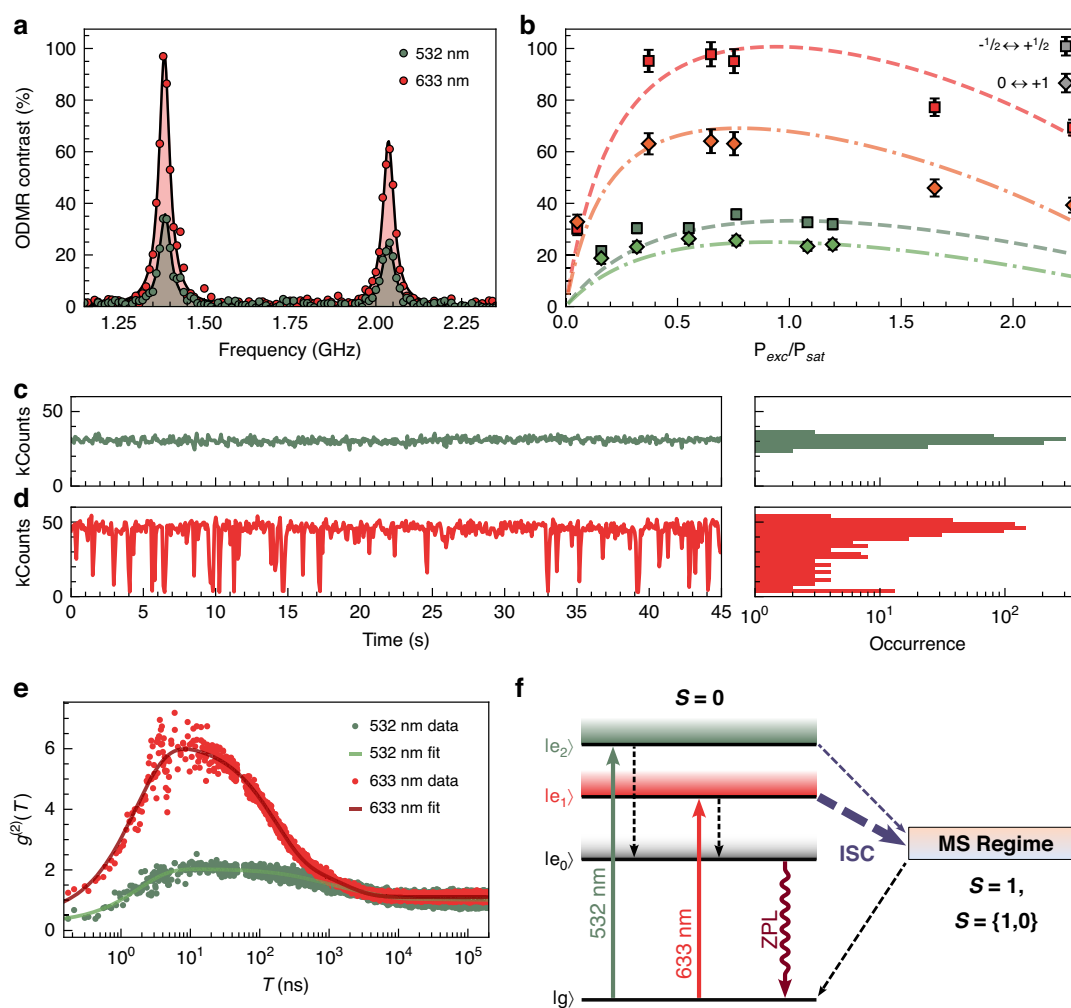


Fig. 2 Behaviour of the spin complex under different excitation wavelengths. **a** CW-ODMR spectra of the emitter excited with 532 nm (green) and 633 nm (red) showing the relative contrast for the $-1/2 \leftrightarrow +1/2$, and $0 \leftrightarrow +1$ (diamonds) resonances under both 532 and 633 nm excitations plotted as a function of optical power normalised by the saturation power ($P_{\text{exc}}/P_{\text{sat}}$). The dashed lines represent the modelled contrast as a function of $P_{\text{exc}}/P_{\text{sat}}$, an explanation of the model can be found in Supplementary Section 4. **c, d** Time traces of the emitter fluorescence under 532 nm and 633 nm excitation, respectively, with the corresponding photon count histograms (1 kCount bin width) shown on the right for each excitation wavelength. **e** Long-timescale second-order autocorrelation measurements, recorded under 532 nm and 633 nm excitation at optical powers of 63 μW and 50 μW , respectively. **f** Energy level diagram illustrating a potential mechanism for population of the MS regime under different excitation wavelengths. Coupling between the state populated by 633 nm excitation and the MS regime is anticipated to be stronger than that of the state populated by 532 nm excitation, leading to increased ODMR contrast and partially responsible for the observed blinking behaviour. Radiative transitions are represented by solid lines, whereas non-radiative transitions are represented by dashed lines

Based on the ODMR contrast results in Fig. 2c, d and the photodynamics observed in Fig. 2c–e, we propose a conceptual model in Fig. 2f that explains the wavelength-dependent behaviour within the current understanding of the energetic structure of the spin complex outlined in Fig. 1b. Different excitation wavelengths, with corresponding photon energies, selectively promote different relaxation pathways within the singlet (optically active, $S = 0$) regime. In Fig. 2f, we present a simplified illustration of the proposed model. Upon absorption of either

633 nm or 532 nm photons, immediate states $|e_1\rangle$ and $|e_2\rangle$ are populated, respectively. The relative energetic positions of these excited states with respect to the $S = 1$ triplet states in the MS regime determine the efficiency of ISC, although it is also possible for them to couple first to an intermediate level that has different decay channels depending on the excitation energy^{40,41}. Considering the simplified model in Fig. 2f, $|e_1\rangle$ is predicted to couple strongly to the MS regime, accounting for the enhanced ODMR contrast and the increased photon bunching seen



in Fig. 2e. In contrast, $|e_2\rangle$ is predicted to have weaker coupling to the MS regime, resulting in less population transfer to the triplet states, thus decreasing ODMR. We applied an existing rate model, used to describe spin complex emitters in hBN, to the power-dependent contrast in Fig. 2b¹³, and observed a close fit to the recorded data. The model assumes a higher ISC rate from $|e_1\rangle$ to the MS regime as compared to $|e_2\rangle$. Additionally, the polarisation rate of the different triplet states following ISC is dependent on the optical pumping power, resulting in the decrease in contrast at higher laser powers as initially by Dréau et al. for NV⁻ centres in diamond⁴² and Patel et al. for single spin defects in hBN⁴³. Details on the derivation and explanation of the model can be found in Supplementary Information Section 4.

To further investigate the excitation wavelength-dependence of the spin complex behaviour discussed above, we extended the same confocal microscope configuration to conduct simultaneous co-excitation with two wavelengths (532 nm and 633 nm). The optical layout is summarised in Fig. 3a as a simplified schematic, and is discussed in detail in ‘Methods’.

Using the co-excitation scheme, we first examine PL dynamics in the time domain. Figure 3b shows PL count traces recorded at a fixed 633 nm excitation power of 80 μ W while varying the 532 nm power from 0 to 150 μ W. As discussed in the previous section, excitation with 633 nm alone produces pronounced blinking. However, introducing 532 nm excitation progressively suppresses this blinking, with higher 532 nm powers leading to increasingly stable emission. This trend is more clearly visualised in the extracted count histograms shown in Fig. 3c, where distributions are plotted on a logarithmic occurrence scale for better visibility.

In addition to suppressed blinking, introduction of 532 nm excitation initially leads to an increase in total PL intensity, followed by counts saturation at higher powers. This is shown in Fig. 3d, where the extracted total PL intensities are plotted as a function of 532 nm power. The dependence can be expressed as $I(P)$, which follows a power saturation model of the form:

$$I(P) = I_{\text{offset}} + \frac{I_{\infty}P}{P + P_{\text{sat}}} \quad (1)$$

The relationship accounts for the constant PL contribution (I_{offset}) from the fixed 633 nm excitation power of 80 μ W and provides luminescence counts I_{∞} at saturation power P_{sat} of 8.33 μ W.

From the histogram plots in Fig. 3c, it is clearly seen that excitation with 633 nm alone results in two distinct count distributions. Using their minimum overlapping point, we define a threshold value (black dashed line), which separates the histograms into ‘dark’ (below

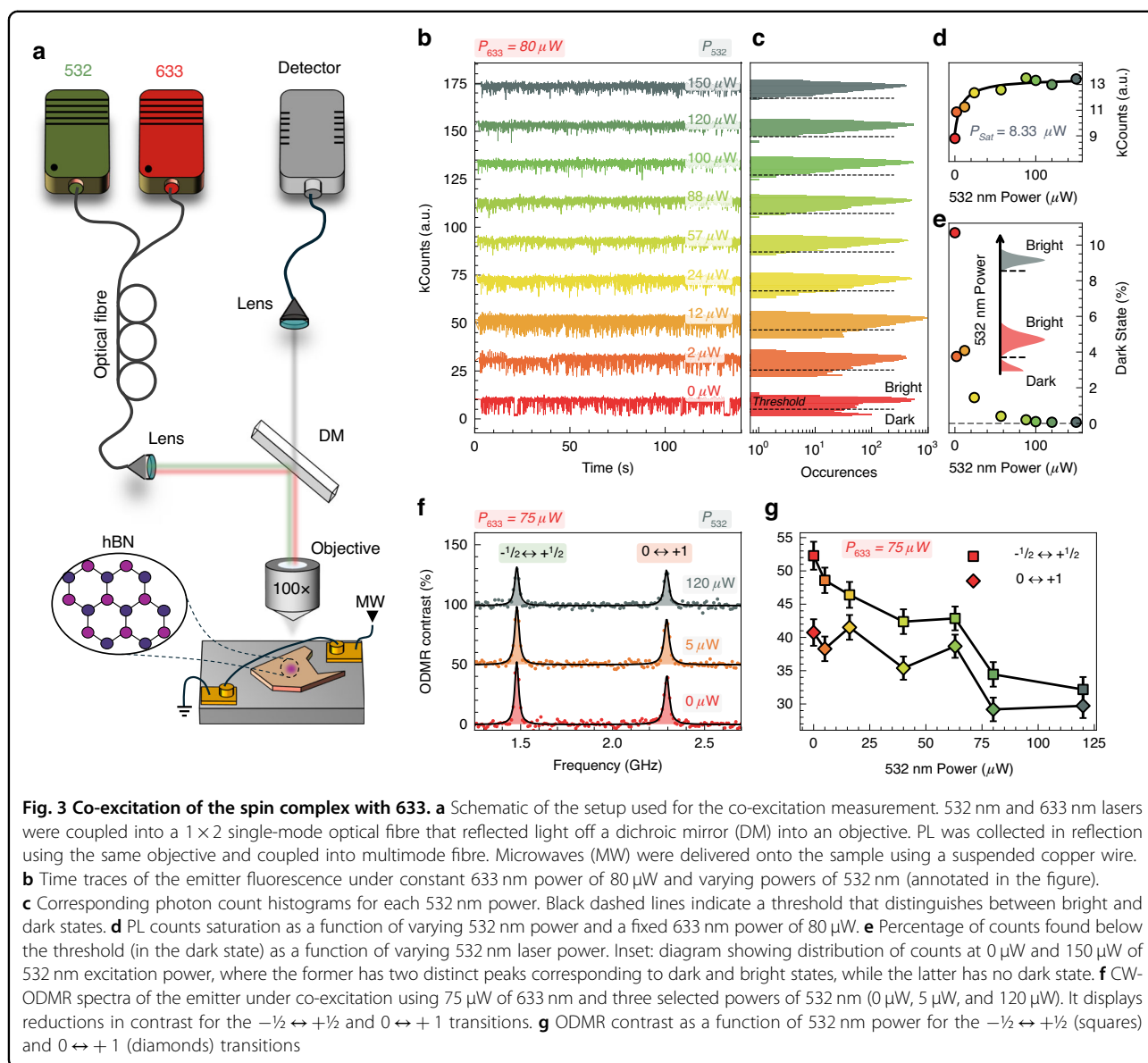
threshold) and ‘bright’ (above threshold) states. The initial threshold $I_{\text{th},0}$ was determined quantitatively from the 633 nm only count histogram as the local minimum between the dark and bright populations, obtained by identifying the two maxima of the histogram and selecting the minimum of the distribution between them. This can be clearly visualised using a simple illustration schematic in the inset of Fig. 3e, where two Gaussian profiles represent two emission states divided by a threshold. For all co-excitation conditions, the threshold value was then scaled according to the relative increase in total PL intensity following the saturation model, $I_{\text{th}}(P) = I_{\text{th},0}I(P)/I(0)$. Using this approach, the density of dark state counts can be consistently extracted at different co-excitation powers of 532 nm (shown in Fig. 3e). Under sole 633 nm excitation, approximately 11% of the counts are in the dark state, whereas addition of only 2 μ W of 532 nm excitation reduces this fraction to ~4%. Increasing the 532 nm power further progressively suppresses the dark state. Equal contributions of 532 nm and 633 nm and dominant 532 nm excitation (≥ 80 μ W) lead to a single bright state above threshold, as shown in Fig. 3c and inset of Fig. 3e.

The observed behaviour suggests that under 633 nm excitation the spin complex more frequently cycles through the MS regime which involves charge transfer between different spin manifolds. Addition of 532 nm excitation may enable optical repumping from the $S = 1$ state into a higher energy excited level⁴⁴. Phonon-assisted decay then allows the electron to relax into a ZPL emissive state that has a reduced coupling to the MS regime. This results in stabilised emission with lower blinking at the expense of ODMR contrast, as electrons are removed from the MS regime.

This interpretation is consistent with the ODMR measurements shown in Fig. 3f, where 633 nm-only excitation yields the highest contrast. Adding low 532 nm power (5 μ W) reduces the contrast by ~4% for $-\frac{1}{2} \leftrightarrow +\frac{1}{2}$ and ~2% for $0 \leftrightarrow +1$ transitions. At the highest 532 nm power, the contrast decreases to ~30% for both transitions. Stepwise measurements as a function of 532 nm power in Fig. 3g reveal a gradual reduction in contrast. However, the different rates at which the dark state population and ODMR contrast decrease suggest that repumping alone may not fully account for the observations. Analysis of other transitions of the $S = 1$ manifold is in Supplementary Section 5.

There are two additional factors contributing to the different trends of the dark state population and ODMR contrast. First, under sole 532 nm excitation the emitter becomes optically active at powers of ≥ 15 μ W. Thus, during co-excitation in the high power regime, increasing dominance of 532 nm excitation competes with the 633 nm pathway, reducing the overall coupling to the MS



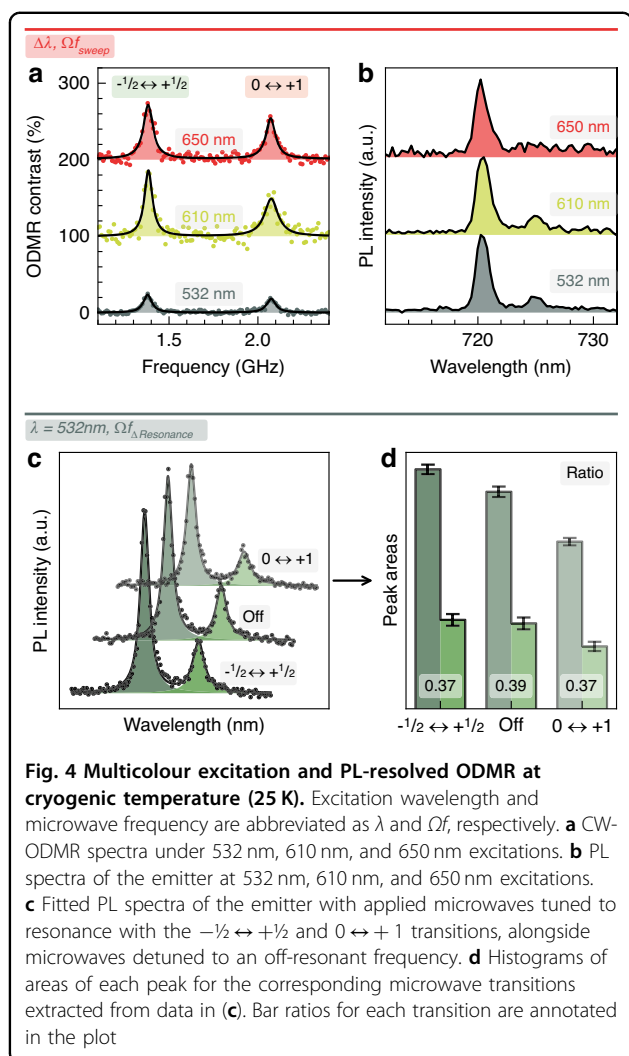


regime. Second, the presence of an alternative non-luminescent decay pathway cannot be excluded. Indeed, it is common for hBN emitters to display blinking characteristics due to having additional charge states^{45,46} or charge traps in the vicinity that are activated through different laser excitation⁴⁷. A combination of these factors contributes to the changes of populations of dark/bright states and ODMR contrast.

In the final section, we probe the spin-complex under multicolour excitation at cryogenic temperatures. Performing these measurements at 25 K allows cleaner spectral resolution by suppressing thermal broadening and reducing background contributions present at room temperature. Figure 4a shows CW-ODMR spectra acquired under 532 nm, 610 nm, and 650 nm excitations.

Using 532 nm, ODMR contrast of the $-1/2 \leftrightarrow +1/2$ and $0 \leftrightarrow +1$ transitions remains below 35%, consistent with room temperature data. Changing the excitation to 610 nm, results in a pronounced increase in ODMR contrast, with both resonances rising above 50%. Changing excitation wavelength even closer to the ZPL (650 nm) yields similar contrast levels, remaining above 50%. The $-1 \leftrightarrow +1$ transition displayed similar behaviour and is shown in Figure SI15. These results highlight once more that for this emitter, excitation wavelengths below ~ 600 nm are less effective at populating the states in the MS regime responsible for high ODMR contrast. Throughout these measurements Fig. 4b confirms that spectral characteristics of the emitter remained unchanged. This then opens avenues for conducting PL spectral response to examine





how microwave-driven spin transitions affect individual emission signatures.

Under 532 nm excitation, the PL spectrum exhibits two dominant peaks separated by ~ 10 meV, consistent with coupling to a longitudinal acoustic phonon mode of the sideband⁴⁸. Figure SI16 shows that both spectral components display minimal fluctuations in intensity and spectral wandering (remaining below ~ 0.3 nm), indicating stable emission from both peaks. Following, we probe the effect of microwave-driven spin transitions on individual spectral components. The applied microwave frequency was set either on resonance with the $-1/2 \leftrightarrow +1/2$ or $0 \leftrightarrow +1$ spin transitions, or to an off-resonant frequency for comparison.

Changes in the relative peak intensities were then analysed. Representative spectra acquired under identical optical conditions with the microwave field tuned on- and off-resonance are shown in Fig. 4c. Double-Lorentzian fits to these spectra yield nearly identical area ratios for the

two peaks across all microwave frequencies. This result is summarised in Fig. 4d, where the extracted peak areas are plotted as histograms and their ratios are annotated. The ratios deviate by less than 0.02, indicating that both spectral components couple evenly with the MS regime. The invariance of the peak-area ratio under the application of microwaves on and off resonance may arise due to the vibronic structure of the ground state manifold. In this case, the dynamics of the excited states (e.g. ISC to the MS regime) occur prior to branching into the two optical channels, resulting in identical changes in both peak areas. This also suggests that the electron relaxation pathways have equal ISC probability to the $S=1$ spin manifold. Thus, the demonstrated technique reliably probes spin dynamics via spectral signatures. We consequently apply this analysis at room temperature to assess the influence of stronger phonon contributions. Similarly, the results revealed minimal variations in peak-area ratios between on- and off-resonant microwave excitation. The corresponding analysis is provided in Supplementary Information Section 7.

The results presented in this work are especially relevant for the development of quantum sensing based on spin complex defects. In particular, the dependence of ODMR contrast on excitation wavelength has direct implications for magnetic field sensing with the spin complex, as the direct current (DC) magnetic field sensitivity (η_{DC}) is calculated as^{42,49}:

$$\eta_{DC} = \mathcal{P}_L \frac{h}{g_e \mu_B} \frac{\Delta\nu}{C\sqrt{R}} \quad (2)$$

Where $\mathcal{P}_L$ is a constant based on the Lorentzian profile of the peak (≈ 0.77), h is Planck's constant, $g_e = 2$ is the electron g-factor, μ_B is the Bohr magneton, $\Delta\nu$ is the ODMR linewidth, C is the contrast, and R is the photon count rate. We find that for the $-1/2 \leftrightarrow +1/2$ transition, η_{DC} improves from $23.1 \pm 2.2 \mu\text{T}\sqrt{\text{Hz}}^{-1}$ under 532 nm excitation to $7.9 \pm 0.7 \mu\text{T}\sqrt{\text{Hz}}^{-1}$ under 633 nm excitation. Similarly, the $0 \leftrightarrow +1$ transition shows sensitivity improvement from $33.1 \pm 4.3 \mu\text{T}\sqrt{\text{Hz}}^{-1}$ to $12.4 \pm 1.5 \mu\text{T}\sqrt{\text{Hz}}^{-1}$ under 532 nm and 633 nm excitations, respectively. The benefit of increased ODMR contrast is accompanied by reduced linewidths under 633 nm excitation compared to 532 nm, resulting in significantly improved sensitivity within the full power range. A detailed analysis is provided in Supplementary Information Section 8, where we also examine the impact of blinking on sensitivity. Specifically, blinking reduces the effective sensing duty cycle and thus the usable photon count rate. To account for this, we introduce an effective count rate that scales with both the intrinsic brightness and the fraction of time spent in the bright state. We find that while blinking introduces measurable changes, the



overall impact on the sensitivity for both transitions is minimal. Notably, the co-excitation scheme provides an optimal balance, where the addition of small 532 nm power ($\sim 16 \mu\text{W}$) suppresses blinking while preserving the high contrast achieved under 633 nm excitation, yielding best sensing performance. Our results establish the excitation wavelength optimisation as a practical pathway for significantly improving contrast readout of the spin complex, useful in quantum sensing applications and applicable to other quantum systems⁵⁰.

Additionally, the wavelength-dependent photodynamics of the spin complex could enhance super-resolution imaging via ground-state depletion microscopy (GSD), as previously demonstrated for hBN emitters⁴⁴. In typical GSD microscopy, doughnut-shaped beam profiles are used to drive emitters surrounding the central intensity null into dark states, thereby increasing the effective resolution at the centre of the null⁵¹. As such, excitation wavelengths that more effectively promote ISC into metastable states are anticipated to enhance the performance of GSD microscopy. This significantly exceeds resolution achieved by a standard confocal microscope that is limited by light diffraction and typically operates within a $\sim 300\text{--}350$ nm range depending on the excitation wavelength. Given the reduction in saturation power under 633 nm excitation, the depletion efficiency is expected to increase accordingly, leading to improved GSD performance compared to 532 nm excitation. As detailed in the Supplementary Information Section 9, this corresponds to an estimated improvement in spatial resolution to 60–65 nm. The spin-dependent fluorescence of the spin complex allows super-resolution imaging to realise quantum sensing at the nanometre scale, using spatial maps of both PL intensity and ODMR contrast.

Discussion

In summary, we performed detailed spin and photodynamics studies of the spin complex emitters in hBN. Remarkably, upon using 633 nm excitation wavelengths, ODMR contrast (and correspondingly the sensitivity) was found to be three times greater than that of the same emitter excited with 532 nm laser. Additionally, the emitter showed different photodynamics, evidenced by strong blinking behaviour and photon bunching present in the autocorrelation measurements. Based on this behaviour, we proposed a conceptual model in which different excitation wavelengths populate distinct excited states with varied coupling strengths to the optically inactive states in the metastable regime. From this, we employed a co-excitation scheme to stabilise PL of the emitter, using relatively weak 532 nm excitation powers in combination with constant 633 nm excitation. Finally, we investigated behaviour of the emitter at cryogenic

temperatures to gain further insight into the spin-dependent PL of the spin complex through spectrally resolved ODMR. Our findings highlight the critical role of excitation wavelength in controlling both the spin-dependent behaviour and photodynamics of spin complex emitters, providing new strategies for optimising their performance in quantum sensing and photonic applications.

Materials and methods

Emitter generation and optical characterisation

Generation of hBN spin-active quantum emitters followed procedures described in our recent work¹². Mechanically exfoliated hBN flakes were annealed in an oxygen environment at 1000 °C for 4 h, following 1 h of UV/ozone treatment. The flakes were then analysed via CW-PL spectroscopy to locate individual emitters.

All optical measurements were carried out using home-built scanning confocal microscope systems operating in reflection measurement. Emitters were addressed using 532 nm and 633 nm lasers coupled into the same 1×2 single-mode optical fibre (Thorlabs) and out-coupled through a fibre launcher equipped with a double-achromatic broadband coated lens. This ensured an identical Gaussian beam profile, diameter, and spatial mode for both excitation wavelengths. Downstream optical components included a red dichroic mirror (Semrock) and an achromatic 100 $\times$ objective (Mitutoyo Plan Apo), typically high numerical aperture (≥ 0.7) to spatially resolve individual emitters. The collected PL signal was coupled into a multimode fibre and detected using either an avalanche photodiode detector (Excelitas APD) or spectrometer charge-coupled detector imaging (Andor SR303I-Newton CCD) using 300 l/mm grating. Correlation measurements used a multimode 50:50 beamsplitter fibre with coincidence counts analysed by a counter module (Swabian Instruments Timetagger 20).

Room temperature measurements were carried out in ambient conditions. For cryogenic measurements, samples were cooled down to 25 K using a closed-loop helium cryostat (Attocube attoDRY 800) and kept at pressure of $\sim 10^{-5}$ mbarr.

Optically detected magnetic resonance

For ODMR measurements, a static magnetic field was applied using a permanent NdFeB magnet (grade N35) positioned perpendicular to the sample surface to lift the spin-state degeneracy. Microwave excitation was provided by a radio-frequency signal generator (AnaPico APSIN 4010), with the output delivered to the sample via a suspended copper wire placed in close proximity to the flake. The microwave signal was amplified using a high-power microwave amplifier (Mini-Circuits ZHL-16W-43-S+). The microwave frequency was swept within the



0.5–4.0 GHz resonance range while monitoring the PL signal to obtain ODMR spectra. The microwave was modulated on and off in alternating acquisition cycles, and the corresponding PL signals were averaged to obtain PL_{on} and PL_{off}, respectively. ODMR contrast was then calculated as $C = (PL_{on} - PL_{off})/PL_{off} \times 100$.

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Author details

¹School of Mathematical and Physical Sciences, University of Technology Sydney, Ultimo, NSW 2007, Australia. ²ARC Centre of Excellence for Transformative Meta-Optical Systems, University of Technology Sydney, Ultimo, NSW 2007, Australia. ³Department of Physics, School of Science, RMIT University, Melbourne, VIC 3001, Australia. ⁴Department of Physics, Kyung Hee University, Seoul 02447, Republic of Korea

Author contributions

I.Z. and N.P.S. contributed equally to the work. M.K. and I.A. conceptualised the project. I.Z., N.P.S., and K.S. performed the experiments. B.W. and J.-P.T. performed the rate modelling and analysed the results. I.Z. and N.P.S. wrote the paper with contribution from all authors. All authors discussed the results and contributed to the paper.

Data availability

The data supporting the findings of this study are available within the paper and its supporting information files.

Conflict of interest

The authors declare no competing interests.

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